Attitude Control of Spacecraft

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Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions



2 INTRODUCTION

- **3** THE KINEMATICS OF A FLEXIBLE SPACECRAFT
- **4** The Dynamics of a Flexible Spacecraft
- **5** THE CONTROL PROBLEM
- **6** SIMULATION RESULTS

7 CONCLUSIONS



2 INTRODUCTION

- **3** THE KINEMATICS OF A FLEXIBLE SPACECRAFT
- THE DYNAMICS OF A FLEXIBLE SPACECRAFT
- **3** THE CONTROL PROBLEM
- **6** SIMULATION RESULTS

7 CONCLUSIONS

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Outline	Introduction ●○○○○	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
Introduction						

Introduction

- In the near future large flexible structures will be used in the space for various uses (ISS, Earth observation, communication, etc.)
- Pointing precision, shape control and integrity of the structures are prior mission requirements
- Large platforms, antennas, solar arrays, etc., of dimensions ranging from some meters possibly up to various kilometers
- To put in orbit these space structures at reasonable costs, their wieght has to be minimized
- It must be possible to compactly store these structures to diminish the costs to put them in orbit (weight and payload space are crucial issues)
- Their dynamic behavior can be difficult to predict analytically (main problems arise from the unreliability or impracticality of structural tests on Earth) and the performances of controls designed on the basis of perfect model knowledge can be deteriorated leading to on-orbit behaviors which can be substantially different from preflight ground test measures or analytic predictions

Outline	Introduction ○●○○○	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
Introduction						

- When the structure dimensions increase, the frequencies of the first natural modes decrease
- The elastic modes of these light structures may be poorly damped, and problem arise if their spectrum overlaps the controller bandwidth
- For these reasons the controller design may be critic and important is to ensure stability and performance

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
Introduction						

- Important control problems
 - 1. Fine alignment with prescribed attitude and precision (payload sensors, antennas, etc. precise pointing)
 - 2. Shape control and integrity of the structures (vibration damping)
 - 3. Great angular displacements for re–orienting the structure (minimum time with minimum fuel consumption)

The control requirements for both the problems are highly demanding

- In attitude control the structure translation is not considered
 - 1. Center of mass translation influence only the spacecraft orbit
 - 2. The spacecraft must stay in a "box"
 - 3. Orbit corrections are performed periodically and when they occur the normal pointing operations are interrupted

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
Introduction						

- The required high control performance can be obtained only considering the rigid–elastic dynamic coupling (nonlinear model)
- Quick review of application of nonlinear control techniques to attitude control
 - RIGID: [Dwyer, IEEE TAC 1984], [Monaco, Stornelli, 1985], [Monaco, Normand-Cyrot, Stornelli, CDC 1986], [Dwyer, CDC 1987], [Wen, Kreutz-Delgado IEEE TAC 1991], [Crouch, IEEE TAC 1984], [Aeyels, S&CL 1985], [Lizarralde, IEEE TAC 1996], [Di Gennaro, Monaco, Normand-Cyrot, Pignatelli, 1997]
 - FLEXIBLE: [Balas, AIAA JG&C 1979], [Joshi, 1989], [Vadali, 1990], [Di Gennaro, CDC 1996], [Di Gennaro, AIAA JGCD 1998], [Di Gennaro, JOTA 1998], etc.
- Important aspect when applying nonlinear controls: state measurement
 - Attitude position and velocity (rigid main body)
 - Modal position and velocity variables (elastic deflection of the flexible appendages)

Outline	Introduction 0000●	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
Introduction						

- Modal position and velocity variables are important for fine pointing and vibration damping – when they are not measured (no appropriate sensors can be used) the control ensuring the performance can not be implemented
- Dynamic nonlinear controllers can reconstruct the modal variables
- Dynamic controllers can also reconstruct main body angular velocity (in case of sensor failure)



2 INTRODUCTION

3 THE KINEMATICS OF A FLEXIBLE SPACECRAFT

THE DYNAMICS OF A FLEXIBLE SPACECRAFT

- **5** THE CONTROL PROBLEM
- **6** SIMULATION RESULTS

7 CONCLUSIONS

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Outline	Introduction	Kinematics ●0000000000	Dynamics	The Control Problem	Simulation Results	Conclusions
Kinematics						

Mathematical Model of a Flexible Spacecraft

• In the derivation of the model we consider first the kinematic equations and then the dynamic ones

Kinematics

• Inertial frame $RC = \{O, x, y, z\}$ and non–inertial frame $R\Gamma = \{\Omega, \xi, \eta, \zeta\}$ (body–fixed frame)



- It is usual to consider $\Omega \equiv O \equiv$ main body center of mass, and $RC \equiv R\Gamma$ at t = 0
- Rr determines the spacecraft attitude (i.e. its position in the space)

Outline	Introduction	Kinematics ○●○○○○○○○○	Dynamics	The Control Problem	Simulation Results	Conclusions
Kinematics						

- When discussing a rotation, there are two possible conventions:
 - Rotation of vectors (the frame remains fixed!)

$$r_{z}^{r_{z}^{\prime}} \xrightarrow{r'} RC$$

$$r_{z} \xrightarrow{\phi} \qquad r' = \bar{\mathcal{R}}(\phi)r, \quad \bar{\mathcal{R}}(\phi) = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \phi & -\sin \phi \\ 0 & \sin \phi & \cos \phi \end{pmatrix}$$

$$x \xrightarrow{r'_{y}} r_{y} \xrightarrow{r_{y}} y$$

 Rotation of a frame RΓ up to coincide with a frame RΓ_r (the vector remains fixed!)



• Note that $\bar{\mathcal{R}}^T = \mathcal{R}$

Outline	Introduction	Kinematics ○○●○○○○○○○	Dynamics	The Control Problem	Simulation Results	Conclusions
Kinematics						

- Considering rotations of $R\Gamma$ of a positive (conterclockwise) angle ϕ
 - About the *x*-axis: The rotation matrix is

$$\mathcal{R}_1(\phi) = egin{pmatrix} 1 & 0 & 0 \ 0 & \cos \phi & \sin \phi \ 0 & -\sin \phi & \cos \phi \end{pmatrix}$$

• About the y-axis:

$$\mathcal{R}_2(\phi) = \begin{pmatrix} \cos\phi & 0 & \sin\phi \\ 0 & 1 & 0 \\ -\sin\phi & 0 & \cos\phi \end{pmatrix}$$

• About the *z*-axis:

$$\mathcal{R}_3(\phi) = egin{pmatrix} \cos \phi & \sin \phi & 0 \ -\sin \phi & \cos \phi & 0 \ 0 & 0 & 1 \end{pmatrix}$$

Outline	Introduction	Kinematics 000●0000000	Dynamics	The Control Problem	Simulation Results	Conclusions
Kinematics						

- Observation: a matrix R(φ) is a linear transformation of vectors (i.e. of the point pointed by r), and can be interpreted as
 - 1. operator which transforms a vector of $R\Gamma$ into a vector of another frame rotated w.r.t. $R\Gamma$.

This vector is *rotated* and we still want to *express it in R*Г.

The angles are positive if *clockwise* (in fact, it is equivalent to rotate the reference frame counterclockwise or rotate the vector clockwise)

- operator which transforms vectors of RΓ into vectors of another frame rotated w.r.t. RΓ. The vectors does not rotate, RΓ rotates (positively if counterclockwise)
- In the following we are interested in the first interpretation. Hence the rotation will be positive if *clockwise*

Outline	Introduction	Kinematics 0000●000000	Dynamics	The Control Problem	Simulation Results	Conclusions
Kinematics						

- Note that $\mathcal{R}_i(\phi)$, i = 1, 2, 3 are
 - Orthogonal ($\mathcal{R}^{-1} = \mathcal{R}^T$)
 - With eigenvalues $\lambda_1 = 1$, $\lambda_{2,3} = e^{\pm j\phi} = \cos \phi \pm j \sin \phi$
 - With determinant equal to 1 (proper orthogonal matrix)
- Transformations represented by such matrices (with unitary eigenvalue) are rotation transformations of an angle ϕ about an axis determined by the corresponding eigenvalue

 $\mathcal{R}\mathbf{v} = \mathbf{v}, \quad \lambda = \mathbf{1}, \quad \mathbf{v} \text{ eigenvector}$

i.e. the axis is the subspace V generated by v, which is transformed into itself during the rotation (namely V is invariant under rotations)

 Euler Theorem Every rotation sequence leaving fixed a point in the space is equivalent to a (certain) rotation of an angle Φ about a (certain) axis passing through this point. The axis is called Euler axis, and is determined by the unit vector ε, and Φ is called Euler angle

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
Kinematics – T	he Euler Angles					

Euler angles

- There exist various parameterizations of a rigid body kinematics
- A minimal set of of parameters allowing the determination of the attitude of a rigid body (spacecraft main body) in the space is that of the Euler angles
- Advantages of Euler angles:
 - have simple geometrical interpretation (rotations about the coordinate axes)
 - can be used to determine the closed form solution of the kinematic equations
 - can be used for spacecraft stabilized about the coordinate axes (small angle maneuvers)
- Disadvantages:
 - kinematics present singularities
 - transcendent functions appear in the kinematics and numeric disadvantages

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
Kinematics – T	he Euler Angles					

- To define the attitude matrix describing the rotation from (inertial) frame *RC* to the spacecraft attitude, i.e. to the (non–inertial) frame $R\Gamma$, one considers a sequence of 3 rotation (each about a coordinate axis) bringing *RC* to superpose on $R\Gamma$
 - $r_1 = \mathcal{R}_i(\varphi)r$

•
$$r_2 = \mathcal{R}_j(\vartheta)r_1 = \mathcal{R}_j(\vartheta)\mathcal{R}_i(\varphi)r$$

•
$$r' = r_3 = \mathcal{R}_k(\psi)r_2 = \mathcal{R}_k(\psi)\mathcal{R}_j(\vartheta)\mathcal{R}_i(\varphi)r$$

• Hence
$$\mathcal{R}_{ijk}(\varphi, \vartheta, \psi) = \mathcal{R}_k(\psi)\mathcal{R}_j(\vartheta)\mathcal{R}_i(\varphi)$$

- *i*, *j*, *k* is the sequence of coordinate axes, about which the rotations are performed
- In total, there are 24 possible sequences

Outline	Introduction	Kinematics ○○○○○○○●○○○	Dynamics	The Control Problem	Simulation Results	Conclusions
Kinematics – The Euler Angles						

• A very used sequence is the 3–1–3 sequence

$$\begin{aligned} \mathcal{R}_{313}(\varphi,\vartheta,\psi) &= \mathcal{R}_3(\psi)\mathcal{R}_1(\vartheta)\mathcal{R}_3(\varphi) \\ &= \begin{pmatrix} \cos\psi & \sin\psi & 0\\ -\sin\psi & \cos\psi & 0\\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & 0 & 0\\ 0 & \cos\vartheta & \sin\vartheta \\ 0 & -\sin\vartheta & \cos\vartheta \end{pmatrix} \begin{pmatrix} \cos\varphi & \sin\varphi & 0\\ -\sin\varphi & \cos\varphi & 0\\ 0 & 0 & 1 \end{pmatrix} \\ &= \begin{pmatrix} \cos\psi\cos\varphi - \cos\vartheta\sin\psi\sin\varphi & \cos\psi\sin\varphi & -\sin\psi\sin\varphi & \sin\vartheta\sin\psi \\ -\sin\psi\cos\varphi - \cos\vartheta\cos\psi\sin\varphi & -\sin\psi\sin\varphi & \sin\vartheta\cos\psi \\ \sin\vartheta\sin\varphi & -\sin\psi\sin\varphi & -\sin\vartheta\cos\varphi & \cos\vartheta \end{pmatrix} \end{aligned}$$

If R₃₁₃(φ, ϑ, ψ) is known (from measurements, etc.) it is possible to calculate the Euler angles (r_{ij} are the elements of R₃₁₃)

$$\vartheta = \arccos r_{33}, \quad \varphi = -\arctan \frac{r_{31}}{r_{32}}, \quad \psi = \arctan \frac{r_{13}}{r_{23}}$$

• Indetermination in ϑ : $\vartheta \in [-\pi, 0)$ or $\vartheta \in [0, \pi)$

Once solved, φ , ψ are uniquely determined

If ϑ multiple of π : only $\varphi + \psi$ (ϑ even multiple), or $\varphi - \psi$ (ϑ odd multiple) are determined (usually sin $\vartheta \ge 0$ or $0 \le \vartheta < \pi$ is considered)

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
	00000	000000000000000000000000000000000000000	000000000000000	000000000	0000000	00
Kinematics – T	he Fuler Angles					

Another popular sequence is the 3–1–2 sequence ("yaw, roll, pitch")

$$\mathcal{R}_{312}(\varphi,\vartheta,\psi) = \begin{pmatrix} \cos\psi\cos\varphi - \sin\vartheta\sin\psi\sin\varphi & \cos\varphi\sin\vartheta\sin\psi & \cos\varphi & -\cos\vartheta\sin\psi \\ -\cos\vartheta\sin\varphi & \cos\vartheta\sin\varphi & \cos\varphi & \sin\vartheta \\ \sin\psi\cos\varphi + \sin\vartheta\cos\psi\sin\varphi & \sin\psi\sin\varphi - \sin\vartheta\cos\psi & \cos\varphi & \cos\varphi \\ \sin\psi\sin\varphi\cos\varphi + \sin\vartheta\cos\psi\sin\varphi & \sin\psi\sin\varphi - \sin\vartheta\cos\psi & \cos\psi & \cos\varphi & \cos\psi \end{pmatrix}$$

- In this case $\vartheta = \arcsin r_{23}$, $\varphi = -\arctan \frac{r_{21}}{r_{22}}$, $\psi = -\arctan \frac{r_{13}}{r_{33}}$
- Indetermination in ϑ : $\vartheta \in [-\pi, -\frac{\pi}{2}) \cup [\frac{\pi}{2}, \pi)$ or $\vartheta \in [-\frac{\pi}{2}, \frac{\pi}{2})$ unless ϑ is an odd multiple of $\frac{\pi}{2}$ – in this case $\vartheta \in [-\frac{\pi}{2}, \frac{\pi}{2})$ (cos $\vartheta \ge 0$)
- For small rotations

$$\mathcal{R}_{ extsf{312}}(arphi,artheta,\psi)\simeq egin{pmatrix} 1&arphi&-\psi\-arphi&1&artheta\\psi&-artheta&1\end{pmatrix}$$



Euler parameters

 Using the Euler axis ε and angle Φ (positive if clockwise) to determine the attitude of the rigid main body



• Vector *r* is transformed into $r' = r + \overrightarrow{PP'} = r + \overrightarrow{PH} + \overrightarrow{HP'}$ with

$$\begin{aligned} Q\dot{P} &= -\vec{\epsilon} \times (\vec{\epsilon} \times O\dot{P}) = -\vec{\epsilon} \times (\vec{\epsilon} \times \vec{v}) \\ \overrightarrow{PH} &= -(1 - \cos \Phi) \overrightarrow{OP} = (1 - \cos \Phi) \vec{\epsilon} \times (\vec{\epsilon} \times r) \\ \overrightarrow{HP'} &= -|\overrightarrow{QP'}| \sin \Phi \frac{\vec{\epsilon} \times \overrightarrow{QP}}{|\overrightarrow{QP}|} = -\sin \Phi \vec{\epsilon} \times \overrightarrow{OP} = -\sin \Phi \vec{\epsilon} \times r \\ &= -\sin \Phi \vec{\epsilon} \times \vec{e} = -\sin \Phi \vec{\epsilon} \times \vec{e} \end{aligned}$$

Outline	Introduction	Kinematics	Dynamics 000000000000	The Control Problem	Simulation Results	Conclusions
Kinematics – T	he Euler Parameters					

• Hence: $r' = r + (1 - \cos \Phi) \epsilon \times (\epsilon \times r) - \sin \Phi \epsilon \times r$

Setting

$$\epsilon \times = \begin{pmatrix} 0 & -\epsilon_3 & \epsilon_2 \\ \epsilon_3 & 0 & -\epsilon_1 \\ -\epsilon_2 & \epsilon_1 & 0 \end{pmatrix} \quad \longleftrightarrow \quad \text{dyadic representation of } \epsilon = \begin{pmatrix} \epsilon_1 \\ \epsilon_2 \\ \epsilon_3 \end{pmatrix}$$

$$\epsilon \times v = \begin{pmatrix} 0 & -\epsilon_3 & \epsilon_2 \\ \epsilon_3 & 0 & -\epsilon_1 \\ -\epsilon_2 & \epsilon_1 & 0 \end{pmatrix} \begin{pmatrix} v_1 \\ v_2 \\ v_3 \end{pmatrix}$$

$$\times (\epsilon \times v) = \begin{pmatrix} 0 & -\epsilon_3 & \epsilon_2 \\ \epsilon_3 & 0 & -\epsilon_1 \\ -\epsilon_2 & \epsilon_1 & 0 \end{pmatrix}^2 \begin{pmatrix} v_1 \\ v_2 \\ v_3 \end{pmatrix} = \begin{pmatrix} -(\epsilon_2^2 + \epsilon_3^2) & \epsilon_1 \epsilon_2 & \epsilon_1 \epsilon_3 \\ \epsilon_1 \epsilon_2 & -(\epsilon_1^2 + \epsilon_3^2) & \epsilon_2 \epsilon_3 \\ \epsilon_1 \epsilon_3 & \epsilon_2 \epsilon_3 & -(\epsilon_1^2 + \epsilon_2^2) \end{pmatrix} \begin{pmatrix} v_1 \\ v_2 \\ v_3 \end{pmatrix}$$
one gets $(\tilde{\epsilon}^2 = \epsilon \epsilon^T - I)$

$$r' = \mathcal{R}(\Phi)r = \begin{bmatrix} I + (1 - \cos \Phi) \tilde{\epsilon}^2 - \sin \Phi \tilde{\epsilon} \end{bmatrix} r = \begin{bmatrix} \cos \Phi I + (1 - \cos \Phi) \epsilon \epsilon^T - \sin \Phi \tilde{\epsilon} \end{bmatrix} r$$

$$\mathcal{R}(\Phi) = \begin{pmatrix} \epsilon_1^2 + (\epsilon_2^2 + \epsilon_3^2) \cos \Phi & \epsilon_1 \epsilon_2 (1 - \cos \Phi) + \epsilon_3 \sin \Phi & \epsilon_1 \epsilon_3 (1 - \cos \Phi) - \epsilon_2 \sin \Phi \\ \epsilon_1 \epsilon_3 (1 - \cos \Phi) + \epsilon_2 \sin \Phi & \epsilon_1 \epsilon_3 (1 - \cos \Phi) - \epsilon_1 \sin \Phi \\ \epsilon_1 \epsilon_3 (1 - \cos \Phi) + \epsilon_2 \sin \Phi & \epsilon_1 \epsilon_3 (1 - \cos \Phi) - \epsilon_1 \sin \Phi \end{pmatrix}$$

 ϵ

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
Kinematics – T	he Euler Parameters					

- This parameterization with the Euler axis uses 3 parameters: Φ and the 3 components of ε (with the constraint ||ε|| = 1)
- Since *ϵ* ∈ = 0, one deduces that *ϵ* is the eigenvector (determines the rotation axis)

$$\mathcal{R}(\Phi)\epsilon = \left[I + (1 - \cos \Phi) \tilde{\epsilon}^2 - \sin \Phi \tilde{\epsilon}\right]\epsilon = \epsilon$$

• ϵ has the same components in RC and in RC

The Unitary Quaternion or Euler symmetric parameters

• Quaternions, introduced by Hamilton in 1843, and used by Whittaker in 1937 to describe the rigid body motion, are defined by

$$q_0 = \cos rac{\Phi}{2}, \quad q = \begin{pmatrix} q_1 \\ q_2 \\ q_3 \end{pmatrix} = \begin{pmatrix} \epsilon_1 \\ \epsilon_2 \\ \epsilon_3 \end{pmatrix} \sin rac{\Phi}{2}$$

- Constraint: $q_0^2 + q_1^2 + q_2^2 + q_3^2 = q_0^2 + \|q\|^2 = 1$
- Since $\sin \Phi = 2\cos \frac{\Phi}{2}\sin \frac{\Phi}{2}, \quad \cos \Phi = 1 - 2\sin^2 \frac{\Phi}{2} \Rightarrow 1 - \cos \Phi = 2\sin^2 \frac{\Phi}{2}$ $\mathcal{R}(\Phi) = \mathcal{R}_{bi} = \begin{pmatrix} \epsilon_1^2 + (\epsilon_2^2 + \epsilon_3^2)\cos \Phi & \epsilon_1\epsilon_2(1 - \cos \Phi) + \epsilon_3\sin \Phi & \epsilon_1\epsilon_3(1 - \cos \Phi) - \epsilon_2\sin \Phi \\ \epsilon_1\epsilon_2(1 - \cos \Phi) - \epsilon_3\sin \Phi & \epsilon_2^2 + (\epsilon_1^2 + \epsilon_3^2)\cos \Phi & \epsilon_1\epsilon_3(1 - \cos \Phi) + \epsilon_1\sin \Phi \\ \epsilon_1\epsilon_3(1 - \cos \Phi) + \epsilon_2\sin \Phi & \epsilon_1\epsilon_3(1 - \cos \Phi) - \epsilon_1\sin \Phi & \epsilon_3^2 + (\epsilon_1^2 + \epsilon_2^2)\cos \Phi \end{pmatrix}$

$$= \begin{pmatrix} q_0^2 + q_1^2 - q_2^2 - q_3^2 & 2(q_1q_2 + q_0q_3) & 2(q_1q_3 - q_0q_2) \\ 2(q_1q_2 - q_0q_3) & q_0^2 - q_1^2 + q_2^2 - q_3^2 & 2(q_2q_3 + q_0q_1) \\ 2(q_1q_3 + q_0q_2) & 2(q_2q_3 - q_0q_1) & q_0^2 - q_1^2 - q_2^2 + q_3^2 \end{pmatrix} = I - 2(q_0I - \tilde{q})\tilde{q}$$

Attitude Control of Spacecraft

- The Control Problem Simulation Results Conclusions Outline Introduction Kinematics **Dynamics** Kinematics - The Unitary Quaternion • Hence $\begin{pmatrix} q_0 \\ a \end{pmatrix}$ describes the spacecraft attitude w.r.t. *RC*, and the transformation matrix describing the rotation that brings RC onto RT
 - is Rhi
 - The guaternion is the generalization of an imaginary number (Hamilton)

$$\mathbf{q} = q_0 + q_1 \mathbf{i} + q_2 \mathbf{j} + q_3 \mathbf{k} = \begin{pmatrix} q_0 \\ q \end{pmatrix}, \qquad q = \begin{pmatrix} q_1 \\ q_2 \\ q_3 \end{pmatrix},$$

where *i*, *j*, *k* are imaginary numbers such that

$$i^{2} = j^{2} = k^{2} = -1$$
 $ij = -ji = k$
 $jk = -kj = i$ $ki = -ik = j$

and q_0 is the real or scalar part and $q = q_1 i + q_2 j + q_3 k$ is the imaginary or vectorial part

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Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
Kinematics – T	he Unitary Quaternior	ı				

- Addition and subtraction are obvious
- Multiplication of 2 quaternions q, e is defined as for complex numbers (but the products of *i*, *j*, *k* are not commutative)

$$\begin{aligned} q_r &= qe = (q_0 + q_1i + q_2j + q_3k)(e_0 + e_1i + e_2j + e_3k) = \\ &= (q_0e_0 - q_1e_1 - q_2e_2 - q_3e_3) + \\ &+ (q_0e_1 + q_1e_0 + q_2e_3 - q_3e_2)i + \\ &+ (q_0e_2 - q_1e_3 + q_2e_0 + q_3e_1)j + \\ &+ (q_0e_3 + q_1e_2 - q_2e_1 + q_3e_0)k = (e_0q_0 - e^Tq, e_0q + q_0e - \tilde{e}q) \end{aligned}$$

i.e.

$$\begin{pmatrix} q_{r0} \\ q_r \end{pmatrix} = \begin{pmatrix} q_0 \\ q \end{pmatrix} \begin{pmatrix} e_0 \\ e \end{pmatrix} = \begin{pmatrix} q_0 e_0 - q_1 e_1 - q_2 e_2 - q_3 e_3 \\ q_0 e_1 + q_1 e_0 + q_2 e_3 - q_3 e_2 \\ q_0 e_2 - q_1 e_3 + q_2 e_0 + q_3 e_1 \\ q_0 e_3 + q_1 e_2 - q_2 e_1 + q_3 e_0 \end{pmatrix}$$
$$= \begin{pmatrix} q_0 & -q^T \\ q & q_0 I + \tilde{q} \end{pmatrix} \begin{pmatrix} e_0 \\ e \end{pmatrix} = \begin{pmatrix} e_0 & -e^T \\ e & e_0 I - \tilde{e} \end{pmatrix} \begin{pmatrix} q_0 \\ q \end{pmatrix}$$

- If q, e represent 2 rotations (defined by the matrices R₁, R₂), the rotation defined by R₂R₁ is represented by the quaternion q_r = qe
- Note the order of the matrices and that of the quaternions

$$\mathcal{R}(q_r) = \mathcal{R}_2(e)\mathcal{R}_1(q) \quad \leftrightarrow \quad \mathrm{q}_r = \mathrm{qe}$$

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
Kinematics – T	he Unitary Quaternior	ı				

Quaternion dynamics

• To derive the quaternion dynamics, let

$$RC \rightsquigarrow R\Gamma_t \rightarrow \boxed{\begin{pmatrix} q_0(t) \\ q(t) \end{pmatrix}} RC \rightsquigarrow R\Gamma_{t+\Delta t} \rightarrow \boxed{\begin{pmatrix} q_0(t+\Delta t) \\ q(t+\Delta t) \end{pmatrix}}$$

the quaternions representing the spacecraft attitude w.r.t. *RC* at time *t* (reference $R\Gamma_t$) and $t + \Delta t (R\Gamma_{t+\Delta t})$

Let

$$R\Gamma_t \rightsquigarrow R\Gamma_{t+\Delta t} \rightarrow \boxed{\begin{pmatrix} e_0 \\ e \end{pmatrix} = \begin{pmatrix} \cos \frac{\Delta \Phi_e}{2} \\ \epsilon_e \sin \frac{\Delta \Phi_e}{2} \end{pmatrix}}$$

the error quaternion representing the spacecraft attitude at time $t + \Delta t$ (i.e. of $(R\Gamma_{t+\Delta t}))$ w.r.t. $R\Gamma_t$

 $\Delta \Phi_e = \Phi(t + \Delta t) - \Phi(t)$ is the rotation performed in Δt about ϵ_e

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Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions				
Kinematics	Kinematics – The Unitary Quaternion									
	/	$\langle \cdot \rangle \rangle$								
	($q_0(t)$	$R\Gamma$	(e_0)						
	(a(4)								
		$q(\iota) / $		LE P						
		$\varphi($	(t)	$\Delta \Psi_e$						
			- ()							
	RC		$\Phi(t+\Delta t)$		Γ					
	n O	($\alpha (+ + \Lambda)$	(+))	$\Delta \mathbf{I} t + \Delta t$					
		(9	$I_0(\iota + \Delta)$	()						
			$g(t+\Delta t)$							
			1 (- '							

For the quaternion multiplication law

$$\begin{pmatrix} q_0(t + \Delta t) \\ q(t + \Delta t) \end{pmatrix} = \begin{pmatrix} e_0 & -e^T \\ e & e_0 I - \tilde{e} \end{pmatrix} \begin{pmatrix} q_0(t) \\ q(t) \end{pmatrix}$$

$$= \begin{pmatrix} \cos \frac{\Delta \Phi_e}{2} & -\epsilon_e^T \sin \frac{\Delta \Phi_e}{2} \\ \epsilon_e \sin \frac{\Delta \Phi_e}{2} & \cos \frac{\Delta \Phi_e}{2} I - \tilde{\epsilon}_e \sin \frac{\Delta \Phi_e}{2} \end{pmatrix} \begin{pmatrix} q_0(t) \\ q(t) \end{pmatrix}$$

Introduction Kinematics

The Control Problem

Simulation Results

Conclusions

Kinematic Equations with the Unitary Quaternions

• As $\Delta t \rightarrow 0$ and for the small angle approximations

Dynamics

$$\cos \frac{d\Phi}{2} \approx 1$$
, $\sin \frac{d\Phi}{2} \approx \frac{d\Phi}{2} = \frac{1}{2} |\omega| dt$

where

Outline

$$\omega(t)| = \lim_{\Delta t \to 0} \frac{\Delta \Phi_e}{\Delta t} = \lim_{\Delta t \to 0} \frac{\Phi(t + \Delta t) - \Phi(t)}{\Delta t} = \frac{d\Phi}{dt} \qquad \omega(t) = \epsilon_e |\omega(t)|$$

Hence

$$\begin{pmatrix} q_0(t+dt)\\ q(t+dt) \end{pmatrix} = \begin{pmatrix} 1 & -\epsilon_{\theta}^{T} \frac{1}{2} |\omega(t)| dt \\ \epsilon_{\theta} \frac{1}{2} |\omega(t)| dt & I - \tilde{\epsilon}_{\theta} \frac{1}{2} |\omega(t)| dt \end{pmatrix} \begin{pmatrix} q_0(t) \\ q(t) \end{pmatrix}$$

$$= \begin{pmatrix} q_0(t) \\ q(t) \end{pmatrix} + \frac{1}{2} \begin{pmatrix} 0 & -\omega^{T}(t) dt \\ \omega(t) dt & -\tilde{\omega}(t) dt \end{pmatrix} \begin{pmatrix} q_0(t) \\ q(t) \end{pmatrix}.$$

$$\Rightarrow \begin{pmatrix} \dot{q}_0(t) \\ \dot{q}(t) \end{pmatrix} = \begin{pmatrix} \frac{q_0(t+dt) - q_0(t)}{dt} \\ \frac{q(t+dt) - q(t)}{dt} \end{pmatrix} = \frac{1}{2} \begin{pmatrix} 0 & -\omega^{T}(t) \\ \omega(t) & -\tilde{\omega}(t) \end{pmatrix} \begin{pmatrix} q_0(t) \\ q(t) \end{pmatrix}.$$

Kinematics The Control Problem Outline Introduction **Dynamics**

Kinematic Equations with the Unitary Quaternions

Rearranging

$$\begin{pmatrix} \dot{q}_0 \\ \dot{q} \end{pmatrix} = \frac{1}{2} \begin{pmatrix} -q^T \\ R(q) \end{pmatrix} \omega = \frac{1}{2} \mathcal{Q}^T(q_0, q) \omega$$

$$egin{aligned} \mathcal{Q}^{ extsf{T}} &= egin{pmatrix} -q ^{ extsf{T}} \ R(q_0,q) &= q_0 T + ilde{q} \ &= egin{pmatrix} q_0 & -q_3 & q_2 \ q_3 & q_0 & -q_1 \ -q_2 & q_1 & q_0 \end{aligned}$$

Note that ω is expressed in $R\Gamma = R\Gamma_t$ Note also that $\omega = 2\mathcal{Q}(q_0, q) \begin{pmatrix} \dot{q}_0 \\ \dot{q} \end{pmatrix}$

- This description is more appropriate for rest-to-rest maneuvers $((q_0, q)$ expresses the attitude error)
- Advantages of (nonminimal) parametrization, with respect to a minimal one (Euler angles)
 - Absence of geometrical singularities
 - 2. Attitude matrix is algebraic and does not depend on transcendental functions
 - Easy quaternion multiplication rule for successive rotations
 - 4. An attitude change is obtained by a single rotation about an appropriate axis (fuel and time optimum)

Kinematic Equations with the Unitary Quaternions

• In the case of attitude *tracking* kinematics are expressed with error quaternion

$$\begin{pmatrix} \dot{\mathbf{e}}_{0} \\ \dot{\mathbf{e}} \end{pmatrix} = \frac{1}{2} \mathcal{Q}^{\mathsf{T}}(\mathbf{e}_{0}, \mathbf{e}) \omega_{\mathbf{e}}$$

$$\omega_{\rm e} = \omega - \omega_{\rm r}$$

• ω_r has to be expressed in $R\Gamma$ (as ω)

$$\omega_r = \mathcal{R}(\boldsymbol{q}_0, \boldsymbol{q}) \mathcal{R}^{\mathsf{T}}(\boldsymbol{q}_{r0}, \boldsymbol{q}_r) \mu(\boldsymbol{q}_r, \dot{\boldsymbol{q}}_r)$$

$$\mathcal{R}(q_0, q) = I - 2(q_0 I - \tilde{q})\tilde{q}$$
$$\mathcal{R}(q_{r0}, q_r) = I - 2(q_{r0} - \tilde{q}_r)\tilde{q}_r$$

with

$$\begin{pmatrix} \dot{q}_{r0} \\ \dot{q}_r \end{pmatrix} = \frac{1}{2} \mathcal{Q}^{\mathsf{T}}(q_{r0}, q_r) \mu(q_r, \dot{q}_r) \quad \Rightarrow \quad \mu(q_r, \dot{q}_r) = 2 \mathcal{Q}(q_{r0}, q_r) \begin{pmatrix} \dot{q}_{r0} \\ \dot{q}_r \end{pmatrix}$$



2 INTRODUCTION

3 The Kinematics of a Flexible Spacecraft

4 The Dynamics of a Flexible Spacecraft

- **5** The Control Problem
- **6** SIMULATION RESULTS

7 CONCLUSIONS

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Outline	Introduction	Kinematics	Dynamics ●●○○○○○○○○○	The Control Problem	Simulation Results	Conclusions
Dynamic Equat	ions					

Dynamic equations

- The dynamic equation for a satellite with flexible elements, such as solar panels, antennas, etc., can be obtained following the hybrid coordinates approach proposed by Likins [Likins, 1970] and used in [Monaco, Stornelli, CDC 1985], [Monaco, Normand-Cyrot, Stornelli, 1986], etc
- Angular velocity dynamics (Euler theorem)

$$\dot{L} = -\omega \times L + u_g + d_r = -(\tilde{\omega}_e + \tilde{\omega}_r)L + u_g + d_r$$

- *L* = total angular momentum (depends on the rigid and flexible dynamics)
- *u*_g = control torque acting (gas jets)
- d_r = disturbances on the main body
- $\omega_e = \omega \omega_r$ (ω_r = reference angular velocity)
- $\omega \times = \tilde{\omega}$ (dyadic representation)

Outline	Introduction	Kinematics	Dynamics ○○●○○○○○○○○	The Control Problem	Simulation Results	Conclusions
Dynamic Equat	tions					

• Reaction wheels dynamics

$$\dot{\Omega} = -\dot{\omega} + J_r^{-1} u_r = -\dot{\omega}_e + J_r^{-1} u_r - \dot{\omega}_r$$

- $\Omega = \mbox{angular velocity of the reaction wheels with respect to the main body$
- *u_r* = reaction wheel driving torques
- J_r = reaction wheel inertia matrix
- Flexible appendage dynamics

$$M_f \ddot{\xi} + C_f \dot{\xi} + K_f \xi = H_f + d_f$$

- $\xi = 3N \times 1$ vector of physical displacements flexible appendages discretized with *N* particles
- M_f , $C_f K_f = 3N \times 3N$ mass, damping and stiffness matrices
- $H_f = 3N \times 1$ vector of noninertial forces due to the main body rotation (centrifugal, Coriolis, due to the non uniform angular velocity variation)
- $d_f = 3N \times 1$ vector of external disturbance on the flexible structure



Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions	
Dynamic Equations							

• Total angular momentum sum of "rigid" and "flexible contributions"

$$L_r = J_t \omega + J_r \Omega$$

$$L_{f} = \sum_{i=1}^{N} (r_{0i} + \xi_{i}) \times m_{i} \dot{\xi}_{i} + L_{f}^{c} = \sum_{i=1}^{N} m_{i} (\tilde{r}_{0i} + \tilde{\xi}_{i}) \dot{\xi}_{i} + \sum_{i=1}^{N} m_{i} (\tilde{r}_{0i}^{T} \tilde{\xi}_{i} + \tilde{\xi}_{i}^{T} \tilde{r}_{0i} + \tilde{\xi}_{i}^{T} \tilde{\xi}_{i}) \omega$$

- m_i = mass of the i^{th} particle ($i = 1, \dots, N$)
- r_{0i} = particle position in the undeformed structure w.r.t. $R\Gamma$
- ξ_i = particle physical displacement
- $J_t = J_{mb} + J_r + \sum_{i=1}^{N} (J_{0i} + m_i \tilde{r}_{0i}^T \tilde{r}_{0i}) =$ inertia matrix of the whole undeformed structure
- J_{mb} = main body inertia matrix
- J_{0i} = particle inertia matrix (expressed in the frame attached to the mass and with the axes parallel to the $R\Gamma$ frame); here $J_{0i} = 0$ since the particles are considered punctiform
- Torsional motion neglected (it can be considered in a similar way)

 Outline
 Introduction
 Kinematics
 Dynamics
 The Control Problem
 Simulation Results
 Conclusions

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 Dynamic Equations
 (Hu)

• Noninertial forces
$$H_f = \begin{pmatrix} H_f \\ \vdots \\ H_{fr} \end{pmatrix}$$
 on the particles

$$H_{fi} = m_i \Big[(\tilde{r}_{0i} + \tilde{\xi}_i) \dot{\omega} - 2\tilde{\omega} \dot{\xi}_i + \tilde{\omega}^T \tilde{\omega} (r_{0i} + \xi_i) \Big] \quad i = 1, \cdots, N$$

- Model for the control law design obtained assuming
 - Small deformations
 - Coriolis and centrifugal effects L^c_f neglected
 - First N_e < N vibration modes (significant modes)

• Hence
$$M_{f0}$$
, C_{f0} , K_{f0} $N_e \times N_e$ matrices and

$$M_{f0}\ddot{\xi}_0 + C_{f0}\dot{\xi}_0 + K_{f0}\xi_0 = H_{f0} + d_{f0}$$

$$\xi_0 = \begin{pmatrix} \xi_{01} \\ \vdots \\ \xi_{0N_e} \end{pmatrix}$$

$$H_f \simeq H_{f0} = \begin{pmatrix} m_1\tilde{r}_{01} \\ \vdots \\ m_{N_e}\tilde{r}_{0N_e} \end{pmatrix} \dot{\omega} = -M_{f0}\delta_0\dot{\omega}$$

$$\delta_0 = \begin{pmatrix} \delta_{01} \\ \vdots \\ \delta_{0N_e} \end{pmatrix} = \begin{pmatrix} \tilde{r}_{01}^T \\ \vdots \\ \tilde{r}_{0N_e}^T \end{pmatrix}$$

$$d_f \simeq d_{f0} = \begin{pmatrix} d_{f01} \\ \vdots \\ d_{f0N_e} \end{pmatrix}$$

$$M_{f0} = \text{diag} \{m_1 I, \cdots, m_{N_e} I\}$$

Attitude Control of Spacecraft

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions	
			00000000000				
Dynamic Equations							

- Decoupling of the flexible dynamics (modal analysis)
 - $\eta = T\xi$ $N_e \times 1$ vector of the appendage modal displacements
 - *T* decouples the dynamics and is orthogonal ($T^{-1} = T^{T}$)

$$TM_{f0}T^{T} = I \qquad TK_{f0}T^{T} = K = \text{ diag } \{\omega_{1}^{2}, \cdots, \omega_{N_{e}}^{2}\}$$

Note that

$$TC_{f0}T^{T} = C \simeq \operatorname{diag} \{2\zeta_{1}\omega_{1}, \cdots, 2\zeta_{N_{e}}\omega_{N_{e}}\}$$

the damping matrix C can be considered diagonal only in first approximation

 λ_i = ω_i² are the first N_e modal frequencies and ζ_i are the damping of the N_e modes

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
Dynamic Equat	tions					

The flexible dynamics are hence

$$\ddot{\eta} + C\dot{\eta} + K\eta = -T\delta_0\dot{\omega} + Td_{f0} = -\delta(\dot{\omega}_e + \dot{\omega}_r) + D_f$$

K, C stiffness and damping matrices, δ the N_e \times 3 coupling matrix

• Also the angular momentum due to the flexible appendages is approximated

$$L_{f} \simeq \sum_{i=1}^{N_{e}} m_{i} \tilde{r}_{0i} \dot{\xi}_{i} = \delta_{0}^{T} M_{f0} \dot{\xi} = \delta_{0}^{T} T^{T} T M_{f0} T^{T} T \dot{\xi} = \delta^{T} \dot{\eta}$$

Hence

$$L = L_r + L_f = J_t(\omega_e + \omega_r) + J_r\Omega + \delta^T \dot{\eta}$$

• The angular velocity dynamics rewrite $\rightarrow \dot{L} = -(\tilde{\omega}_e + \tilde{\omega}_r)L + u_g + d_r$

$$J_t(\dot{\omega}_{e}+\dot{\omega}_{r})+J_r\dot{\Omega}+\delta^T\ddot{\eta}=-(\tilde{\omega}_{e}+\tilde{\omega}_{r})\bigg[J_t(\omega_{e}+\omega_{r})+J_r\Omega+\delta^T\dot{\eta}\bigg]+u_g+d_r$$

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions		
Mathematic	al Model							
• 1	The mathen	natical mode	el of the flex	xible spacecra	ft is			
$\dot{e}_0 =$	$-\frac{1}{2}e^{T}\omega_{e} \leftarrow$	— redundant		Green = ki	nematics			
$\dot{e} = \frac{1}{2}R(e)\omega_e$				Blue = rigid dynamics				
$\dot{\omega}_{e}=$	$J_{mb}^{-1} \Big[- (ilde{\omega}_e +$	$-\tilde{\omega}_r)(J_t\omega_e+v)$	$J_r\Omega + \delta^T z +$	$J_t\omega_r) + \delta^T (Cz -$	$+K\eta)+u_g-u_r$	$+ D_r \Big] - \dot{\omega}_r$		
$\dot{\Omega} =$	$-\dot{\omega}_e + J_r^{-1} u_r$	$-\dot{\omega}_r$						
$\dot{\eta} =$	Z			Red = flexibl	e dynamics			
ż=	$-\delta\dot{\omega}_{e} - (Cz$	$+ K\eta) - \delta \dot{\omega}_r$	$+ D_f$					

with
$$D_r = d_r - \delta^T D_f = d_r - \delta_0^T d_{f0}$$
, $J_t = J_{mb} + J_r + \delta^T \delta$

• Sometime it is practical to substitute $\dot{\Omega}$ with

$$\dot{L} = -(\tilde{\omega}_e + \tilde{\omega}_r)L + u_g + d_r$$

 The disturbances acting on the structure are supposed negligible for simplicity

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
Mathematical M	lodel					

• In order to simplify the design of the control law, we use the variables

$$\begin{aligned} \xi_{e} &= L - J_{t}\omega_{r} = J_{t}\omega_{e} + J_{r}\Omega + \delta^{T}\dot{\eta} = (J_{mb} + J_{r})\omega_{e} + J_{r}\Omega + \delta^{T}\psi \\ \psi &= \delta\omega_{e} + \dot{\eta} \end{aligned}$$

- ξ_e = error between total angular momentum and reference angular momentum of the undeformed structure with idle reaction wheels
- ψ = difference between the total modal velocity $\delta \omega + \dot{\eta}$ and the reference velocity $\delta \omega_r$, in modal coordinates
- Therefore, the *mathematical model of a flexible spacecraft* can be rewritten

$$\dot{e}_{0} = -\frac{1}{2}e^{T}\omega_{e} \quad \longleftarrow \text{ redundant}$$

$$\dot{e} = \frac{1}{2}R(e)\omega_{e} \qquad \underset{1}{\overset{N(\omega_{e}, \xi_{e}, \omega_{r}) = (\tilde{\omega}_{e} + \tilde{\omega}_{r})(\xi_{e} + J\omega_{r})}}{\dot{\omega}_{e} = J_{mb}^{-1}\left[-N(\omega_{e}, \xi_{e}, \omega_{r}) + \delta^{T}\left(C\psi + K\eta - C\delta\omega_{e}\right) + u_{g} - u_{r}\right] - \dot{\omega}_{r}}$$

$$\dot{\xi}_{e} = -N(\omega_{e}, \xi_{e}, \omega_{r}) + u_{g} - J\dot{\omega}_{r}$$

$$\begin{pmatrix}\dot{\eta}\\\psi\end{pmatrix} = A\begin{pmatrix}\eta\\\psi\end{pmatrix} - AB\delta\omega_{e} - B\delta\dot{\omega}_{r} \qquad \leftarrow A = \begin{pmatrix}0 & I\\-K & -C\end{pmatrix}, \quad B = \begin{pmatrix}0\\0\\\psi\end{pmatrix} = A = \begin{pmatrix}0\\\psi\end{pmatrix} = A(\xi_{e}, \xi_{e}, \xi_{e}) = A(\xi_{e}, \xi_{e})$$

Attitude Control of Spacecraft

Outline	Introduction	Kinematics	Dynamics ○○○○○○○○○●	The Control Problem	Simulation Results	Conclusions
Mathematical M	lodel					

• When dealing with *rest-to-rest maneuver* (i.e. $q_r = 0, \omega_r = 0, \dot{\omega}_r = 0$)

$$\begin{split} \dot{q}_{0} &= -\frac{1}{2}q^{T}\omega \\ \dot{q} &= \frac{1}{2}R(\boldsymbol{e})\omega \\ \dot{\omega} &= J_{mb}^{-1}\Big[-N(\omega,\xi) + \delta^{T}\Big(C\psi + K\eta - C\delta\omega\Big) + u_{g} - u_{r}\Big] \\ \dot{\xi} &= -N(\omega,\xi) + u_{g} \\ \begin{pmatrix} \dot{\eta} \\ \dot{\psi} \end{pmatrix} &= A\begin{pmatrix} \eta \\ \psi \end{pmatrix} - AB\delta\omega \end{split}$$

with

$$\xi = \mathbf{L} = \mathbf{J}_t \omega + \mathbf{J}_r \Omega + \delta^T \dot{\eta} = (\mathbf{J}_{mb} + \mathbf{J}_r) \omega + \mathbf{J}_r \Omega + \delta^T \psi \qquad \qquad \mathbf{N}(\omega, \xi) = \tilde{\omega} \xi$$



2 INTRODUCTION

- **3** THE KINEMATICS OF A FLEXIBLE SPACECRAFT
- THE DYNAMICS OF A FLEXIBLE SPACECRAFT
- **5** THE CONTROL PROBLEM
- **6** SIMULATION RESULTS

7 CONCLUSIONS

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Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
The Control Pro	oblems					

Problem Formulation

 Rest-to-rest maneuvers: Drive RΓ to RΓ_r (constant), damping out the induced flexible oscillations

$$\lim_{t\to\infty} q = 0 \qquad \lim_{t\to\infty} \eta = 0$$

 Tracking maneuvers: RΓ tracks RΓ_r (variable), damping out the induced flexible oscillations

$$\lim_{t\to\infty} \mathbf{e} = \mathbf{0}, \qquad \lim_{t\to\infty} \eta = \mathbf{0}$$

• Problems solved for $\sigma(A) \in \mathbb{C}^-$ (non–negligible internal damping)

Kinematics The Control Problem Outline Introduction **Dynamics** 00000000

Simulation Results

 $\dot{q}_0 = -rac{1}{2}q^T\omega$

 $\dot{q} = \frac{1}{2}R(e)\omega$

Conclusions

The Rest-to-Rest Maneuver

Rest-to-Rest Maneuvers

State/Output–Feedback Controllers

- Simple case: rigid spacecraft →
- $\dot{\omega} = J_{mb}^{-1} \Big[-\omega \times J_{mb} \omega + u_g \Big]$ State Feedback. A simple proportional and derivative control $u_{\alpha} = -k_{\alpha}q - k_{d}\omega$, with $k_p > 0$, $k_d > 0$ scalars, globally asymptotically stabilizes a rigid spacecraft [Wie, AIAA JG 1985]
- In fact, take the following Lyapunov function candidate

$$V = k_p \Big[(q_0 - 1)^2 + q^T q \Big] + \frac{1}{2} \omega^T J_{mb} \omega$$
$$\downarrow$$
$$\dot{V} = k_p q^T \omega + \omega^T \Big[-\omega \times J_{mb} \omega - k_p q - k_d \omega \Big] = -k_d \omega^T \omega \le 0$$

La Salle theorem:

$$\mathbf{x}(t) \rightarrow \{\mathbf{q} = \mathbf{0}, \omega = \mathbf{0}\} = \mathcal{E} \subseteq \mathbf{E} = \{\mathbf{x} \in \mathbf{R}^n \mid \dot{\mathbf{V}} = \mathbf{0}\}$$



Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
The Rest-to-Re	est Maneuver					

- Output Feedback. Modal position and velocity are often not measurable (sensors placed along the flexible structure necessary)
- Extend the previous controller: estimates $\hat{\eta}, \hat{\psi}$ of η, ψ
- Lyapunov function candidate $P_2 = P_2^T > 0$

• The control law is now slightly changed

$$u = -k_{p}q - k_{d}\omega - \delta^{T} \left[\begin{pmatrix} K \\ C \end{pmatrix} - P_{1}AB \right] \begin{pmatrix} \hat{\eta} \\ \hat{\psi} \end{pmatrix}$$

and the update law is chosen

$$\begin{pmatrix} \hat{\eta} \\ \hat{\psi} \end{pmatrix} = A \begin{pmatrix} \hat{\eta} \\ \hat{\psi} \end{pmatrix} - AB\delta\omega + P_2^{-1} \left[\begin{pmatrix} K \\ C \end{pmatrix} - P_1 AB \right] \delta\omega$$

$$\Rightarrow \dot{V} = -\omega^T (k_d I + \delta^T C \delta) \omega - \begin{pmatrix} \eta^T & \psi^T \end{pmatrix} Q_1 \begin{pmatrix} \eta \\ \psi \end{pmatrix} - \begin{pmatrix} e_\eta \\ e_\psi \end{pmatrix} Q_2 \begin{pmatrix} e_\eta \\ e_\psi \end{pmatrix} \le 0$$

$$\Rightarrow \text{ La Salle} \Rightarrow \text{ global stability}$$

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
The Rest-to-R	est Maneuver					

Tracking Maneuvers State/Output–Feedback Controllers

- La Salle theorem can not be applied Barbalat theorem instead
- State Feedback. Take the PD-like controller

$$u_g = -k_p e - k_d \omega_e - \frac{1}{2} J_{mb} R(e) \omega_e + N - \delta^T (C\psi + K\eta - C\delta\omega_e) + J_{mb} \dot{\omega}_r$$

designed using

$$V(t, x) = (k_{\rho} + k_{d}) \Big[(e_{0} - 1)^{2} + e^{T} e \Big] + \frac{1}{2} (e + \omega_{e})^{T} J_{mb}(e + \omega_{e}) + \frac{1}{2} \Big(\eta^{T} - \psi^{T} \Big) P \begin{pmatrix} \eta \\ \psi \end{pmatrix}$$

Deriving

$$\begin{split} \dot{V}(t,x) &= (k_{p} + k_{d})e^{T}\omega_{e} + (e + \omega_{e})^{T} \left[\frac{1}{2}J_{mb}R(e)\omega_{e} - \bar{N}(\omega_{e},\psi,\omega_{r}) + \delta^{T}(C\psi + K\eta - C\delta\omega_{e}) + u_{g} - J_{mb}\dot{\omega}_{r}\right] + \left(\eta^{T} \quad \psi^{T}\right)P\left[A\begin{pmatrix}\eta\\\psi\end{pmatrix} - AB\delta\omega_{e} - B\delta\dot{\omega}_{r}\right] \\ &\stackrel{u_{g}}{=} - k_{p}||e||^{2} - k_{d}||\omega_{e}||^{2} - \left\|\frac{\eta}{\psi}\right\|_{Q}^{2} - \left(\eta^{T} \quad \psi^{T}\right)PAB\delta\omega_{e} - \left(\eta^{T} \quad \psi^{T}\right)PB\delta\dot{\omega}_{r} \\ &\leq -\lambda_{m}||x||^{2} + \alpha||\dot{\omega}_{r}|||x|| \\ \bullet \text{ Hence } \|x\| \to 0 \text{ (Barbalat)} \end{split}$$

Outline	Introduction	Kinematics	Dynamics	The Control Problem ○○○○○●○○○	Simulation Results	Conclusions
The Rest-	o-Rest Maneuver					
9 9	<i>Proof.</i> We supp Integrating both	ose that $\omega_r \in$ sides of $\dot{V}(t)$	$\in L_\infty[0,\infty)$ and $f, oldsymbol{x}) - \lambda_m \ oldsymbol{x}\ ^2$	d $\dot{\omega}_r \in L_2[0,\infty) \cap$ + $\alpha \ \dot{\omega}_r\ \ x\ $ we h	$L_\infty[0,\infty)$ ave	
١	$V(t, x) - V(0, x_0)$	$(0) \leq -\lambda_m \int_0^\infty$	$\ \mathbf{x}(\tau)\ ^2 d\tau +$	$\alpha \int_0^t \ \dot{\omega}_r(\tau)\ \ \mathbf{x}(\tau)\ $	$)\ { extsf{d}} au \leq$	
Schwar	z inequality —	$\rightarrow \leq -\lambda_m \int_0^{\infty}$	$\ \mathbf{x}(au)\ ^2 d au +$	$\alpha \left[\int_0^t \ \dot{\omega}_r(\tau)\ ^2 d\tau \right]$	$\int_{0}^{1/2} \left[\int_{0}^{t} \ x(\tau)\ ^{2} \right]$	$d au ight]^{1/2}$
	and considering	the limit as	t tends to infini	ty ($\ \cdot\ _2$ is the L_2 -	norm) one has	
		$V(\infty, x)$	$-V(0, x_0) \leq -$	$-\lambda_m \ \mathbf{x}\ _2^2 + \alpha \ \dot{\omega}_r\ _2^2$	$ x _2$	
۹	Since $V(\infty, x)$	\geq 0,				
	λ_{i}	$\ x\ _{2}^{2} - \alpha \ x\ _{2}^{2}$	$\dot{v}_r\ _2 \ \mathbf{x}\ _2 \leq V($	$(0, x_0) - V(\infty, x)$	$\leq V(0, x_0)$	
	and because $\dot{\omega}_r$	$r \in L_2[0,\infty)$	\cap $L_\infty[0,\infty)$ we	e obtain the bound	l	
	$\ x\ _2 \le \frac{1}{\sqrt{\lambda}}$	$\underline{=} \left[V(0, x_0) \right]$	$+ \frac{\alpha^2}{4\lambda_m} \ \dot{\omega}_r\ _2^2 \bigg]$	$^{1/2} + \frac{\alpha}{2\lambda_m} \ \dot{\omega}_r\ _2$	$\Rightarrow x \in L_2[0,\infty)$	ວ)
•	It follows that V trajectories and Also \dot{x} is uniform	$(\infty, x) < \infty$, therefore, x nly bounded	, i.e. $V(t, x)$ is is uniformly be and, hence, x	uniformly bounded bunded is uniformly contin	I in t along the solution uous. Since x is a	ıtion
	uniformly contin	uous functio	n in $L_2[0,\infty)$, v	we have that (Barb	alat theorem)	
	$\lim_{t \to \infty}$	° x = 0 a	and in particula	$\operatorname{r} \lim_{t\to\infty} e = 0,$	$\lim_{t\to\infty}\eta=0$	_

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
The Rest-to-Re	est Maneuver					

Remark. The control u_{PD} is sufficient to solve the attitude tracking when ω_r, ώ_r ∈ L₂[0,∞) ∩ L_∞[0,∞). In fact, taking

$$V(t, \mathbf{x}) = (k_{p} + k_{d}) \Big[(e_{0} - 1)^{2} + e^{T} e \Big] + \frac{1}{2} (e^{T} + \omega_{e}^{T}) J_{mb}(e + \omega_{e}) \\ + \frac{1}{2} \Big(\eta^{T} \quad \psi^{T} \Big) \Big[P + \begin{pmatrix} \mathbf{K} & \mathbf{0} \\ \mathbf{0} & \mathbf{I} \end{pmatrix} \Big] \begin{pmatrix} \eta \\ \psi \end{pmatrix}$$

finally one works out (α_1 , α_2 are appropriate constants)

$$\dot{V}(t, \mathbf{x}) \leq -\lambda_m \|\mathbf{x}\|^2 + (\alpha_1 \|\omega_r\| + \alpha_2 \|\dot{\omega}_r\|) \|\mathbf{x}\| \quad \Rightarrow \quad \lim_{t \to \infty} \mathbf{x} = \mathbf{0}$$

- The PD controller is robust in the sense that it does not need the parameter knowledge
- Moreover, can be used when measurements of the modal variables are not available
- Nevertheless, the gains k_p , k_d must be higher (no good in space)

Outline	Introduction	Kinematics	Dynamics	The Control Problem ○○○○○○●○	Simulation Results	Conclusions
The Rest-to-Re	est Maneuver					

• Output Feedback. If the modal variables η , ψ are not measured, use the same control but with estimates $\hat{\eta}$, $\hat{\psi}$

$$u_{g} = -k_{p}e - k_{d}\omega_{e} - \frac{1}{2}J_{mb}R(e)\omega_{e} + \hat{N} - \delta^{T}(C\hat{\psi} + K\hat{\eta} - C\delta\omega_{e}) + J_{mb}\dot{\omega}_{r}$$

• The design of the update controls $\dot{\hat{\eta}},\, \hat{\psi}$ is based on

$$\mathcal{V}(t, x, e_{\eta}, e_{\psi}) = V(t, x) + rac{1}{2} \begin{pmatrix} e_{\eta}^{\intercal} & e_{\psi}^{\intercal} \end{pmatrix} \Gamma^{-1} \begin{pmatrix} e_{\eta} \\ e_{\psi} \end{pmatrix}$$

Analogously to the previous case one eventually gets

$$\begin{pmatrix} \dot{\hat{\eta}} \\ \dot{\hat{\psi}} \end{pmatrix} = \mathbf{A} \begin{pmatrix} \hat{\eta} \\ \hat{\psi} \end{pmatrix} - \mathbf{A} \mathbf{B} \delta \omega_{\mathbf{e}} - \mathbf{B} \delta \dot{\omega}_{\mathbf{r}} + \Gamma \begin{pmatrix} \mathbf{K} \delta \\ \delta (\tilde{\omega}_{\mathbf{e}} + \tilde{\omega}_{\mathbf{r}}) + \mathbf{C} \delta \end{pmatrix} (\mathbf{e} + \omega_{\mathbf{e}})$$

with

$$\hat{N} = (\tilde{\omega}_{e} + \tilde{\omega}_{r})(J_{mb}\omega_{e} + \delta^{T}\hat{\psi} + J_{s}\omega_{r})$$

Outline	Introduction	Kinematics	Dynamics	The Control Problem ○○○○○○○●	Simulation Results	Conclusions
The Rest-to-	Rest Maneuver					
Rea When Q Ly (/	tusing react yapunov $V(t, x)$ $(k_{\rho} + k_{d}) [(e_{0} - t_{\rho})]$	ion wheels $) = 1)^{2} + e^{T}e + \frac{1}{2}$ iately	the metho $\frac{1}{2}(e+\omega_e)^T J_n$	ds substantial $_{nb}(e+\omega_e) + \frac{1}{2}(\eta_{nb})$	ly is the same ${}^{T} \psi^{T} P \begin{pmatrix} \eta \\ \psi \end{pmatrix} + \frac{1}{2}$.ξ <mark>.</mark> ξε
u N	$r = -\left[-k_{p}e - \left[-k_{p}e - \tilde{\omega}_{r}\right](\varepsilon)\right]$	$-k_d\omega_e - \frac{1}{2}J_{mb}$ $(e_e + J\omega_r)$	$R(e)\omega_e + N$	$-\delta'(C\psi+K\eta-$	$C\delta\omega_e)+J_{mb}\dot{\omega}_r]+\xi$	$\xi_e J \omega_r$

Solution Calculate
$$\dot{V}(t, \mathbf{x}) \leq -\lambda_m \|\mathbf{x}\|^2 + \alpha_0 \varphi(\omega_r, \dot{\omega}_r) \|\mathbf{x}\|$$

() Under stronger hypothesis $\omega_r, \dot{\omega}_r \in L_2[0,\infty) \cap L_\infty[0,\infty)$

Sarbalat $\lim_{t\to\infty} x = 0$, while $\lim_{t\to\infty} \Omega = J_r^{-1} J\omega_r$ (obvious)

When
$$\eta$$
, ψ are not measured use estimates

$$u_{r} = -\left[-k_{\rho}e - k_{d}\omega_{e} - \frac{1}{2}J_{mb}R(e)\omega_{e} + \hat{N} - \delta^{T}(C\hat{\psi} + K\hat{\eta} - C\delta\omega_{e}) + J_{mb}\dot{\omega}_{r}\right] + \tilde{\xi}_{e}J\omega_{r}$$

$$\hat{N} = N(\omega_{e}, \hat{\xi}_{e}, \omega_{r}) = (\tilde{\omega}_{e} + \tilde{\omega}_{r})(\hat{\xi}_{e} + J\omega_{r})\hat{\xi}_{e} = (J_{mb} + J_{r})\omega_{e} + J_{r}\Omega + \delta^{T}\hat{\psi}$$

② Prove stability using $\mathcal{V}(t, x, e_{\eta}, e_{\psi}) = V(t, x) + \frac{1}{2}\xi_{e}^{T}\xi_{e} + \frac{1}{2} \begin{pmatrix} e_{\eta}^{T} & e_{\psi}^{T} \end{pmatrix} \Gamma^{-1} \begin{pmatrix} e_{\eta} \\ e_{\psi} \end{pmatrix}$

• Get
$$\begin{pmatrix} \hat{\eta} \\ \hat{\psi} \end{pmatrix} = A \begin{pmatrix} \hat{\eta} \\ \hat{\psi} \end{pmatrix} - AB\delta\omega_e - B\delta\dot{\omega}_r + \Gamma \begin{pmatrix} K\delta \\ \delta(\tilde{\omega}_e + \tilde{\omega}_r) + C\delta \end{pmatrix} (e + \omega_e)$$

• Barbalat again



2 INTRODUCTION

- **3** THE KINEMATICS OF A FLEXIBLE SPACECRAFT
- THE DYNAMICS OF A FLEXIBLE SPACECRAFT
- **3** THE CONTROL PROBLEM
- **6** SIMULATION RESULTS

7 CONCLUSIONS

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Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
Simulation Res	ults					

Simulation Results (Tracking Maneuvers)

- The mathematical model of the spacecraft has been implemented on a digital computer
- Spillover has been studied by considering a model with more elastic modes than in the model used to derive the control law

Material	Aluminum	Shear modulus	2.5 10 ¹⁰ N/m ²
Length	20 m	r _x	1.5 m
Density	2.76 10 ³ Kg/m ³	r _y	2.3 m
Young modulus	6.8 10 ¹⁰ N/m ²	٢z	-0.8 m

Characteristics of the flexible appendage

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results ○●○○○○○	Conclusions
Simulation Res	ults					

- Three elastic modes result from the modal analysis of the structure, with natural frequencies $\omega_{n1} = 19.38$, $\omega_{n2} = 77.98$, $\omega_{n3} = 157.22$ rad/s and dampings $\zeta_1 = 0.0001$, $\zeta_2 = 0.00005$, $\zeta_3 = 0.00001$
- Only the first two modes have been considered in the controller design
- Coupling matrix δ

$$\delta = \begin{pmatrix} 14.3961 & 8.37634 & -5.29354 \\ -20.4871 & 7.59188 & -6.08014 \\ 4.50401 & 11.5222 & -12.6033 \end{pmatrix} \text{ Kg}^{1/2}\text{m}$$

- A payload of 30 Kg is present at the tip of the appendage
- The spacecraft is characterized by an inertia matrix

$$J_{mb} = \begin{pmatrix} 400 & 3 & 10 \\ 3 & 300 & 12 \\ 10 & 12 & 200 \end{pmatrix} \text{ Kg m}^2$$

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
Simulation Res	ults					

Reference trajectory described as

$$q_{r0} = \cos\frac{\phi_r}{2}, \quad q_r = \begin{pmatrix} \cos 0.5 \ t \\ \sin 0.5 \ t \\ 0 \end{pmatrix} \sin\frac{\phi_r}{2}, \quad \phi_r = \sin\gamma t, \quad \gamma = 0.035 \text{ rad/s}$$

- Trajectory corresponding to a spiral maneuver which, starting from the initial spacecraft attitude, diverges when ϕ_r increases and converges when ϕ_r decreases
- $R\Gamma \equiv R\Gamma_r$ for t = 0, i.e. $e_0(0) = 1$, $e_i(0) = 0$ i = 1, 2, 3; the initial error angular velocity is $\omega_e(0) = \omega_r(0)$
- $\eta(0) = 0, \psi(0) = \delta \omega_e(0) + \dot{\eta}(0) = \delta \omega_r(0)$ (undeformed flexible appendages)

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
Simulation Res	ults					

- A comparison between the controllers with gas gets (state feedback and output feedback) has been conducted
- For both the controllers $k_p = 10^5$, $k_d = 3 \times 10^5$
- The simulations are rendered more realistic by respecting the fact that the gas jets work in a "bang–bang" manner, with saturation values at 60 Nm. This renders harder the control task



 A PD controller is capable to track the desired trajectory when the angular velocity is low, but when it increases and the influence of the flexibility becomes too high and unstable, the input saturation



PD Controller - Actual (solid) and reference (dotted) quaternion components

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
Simulation Res	sults					

• The state feedback controller is capable to track the reference. The control effort (norm of *u_g*) similar to the PD case



State Controller – Actual (solid) and reference (dotted) quaternion components

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions
Simulation Res	ults					

• Analogous result for the output feedback controller. η and ψ are well estimated



Output Controller - Actual (solid) and reference (dotted) quaternion components

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions



2 INTRODUCTION

- **3** THE KINEMATICS OF A FLEXIBLE SPACECRAFT
- THE DYNAMICS OF A FLEXIBLE SPACECRAFT
- **3** THE CONTROL PROBLEM
- **6** SIMULATION RESULTS

7 CONCLUSIONS

Conclusions

- Control laws for rigid/flexible spacecraft can be determined with the Lyapunov approach
- When elastic variables are not known, dynamics in the controller ensure stability
- The absence of measurements of the modal variables is a clear advantage for practical implementations
- $\bullet\,$ The method can be extended to the case of ω not measured
- Extensions are also possible with estimation of perturbations
- The knowledge of system parameters, in particular those describing the elastic motion (natural frequencies and damping ratios), is an obvious limitation, since they are not usually known accurately
- Adaptive robust controllers can avoid this drawback

Outline	Introduction	Kinematics	Dynamics	The Control Problem	Simulation Results	Conclusions ○●
Conclusions						

Thank You!

See details in

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